

Tutorial sheet 8

Discussion topics: What are the fundamental equations of the dynamics of a relativistic fluid? What is the relation between the energy-momentum tensor of a perfect relativistic fluid and its internal energy, pressure, and four-velocity? How is the latter defined?

23. Propagation of internal waves in the ocean

This exercise is inspired by an article (in French) by L. Gostiaux and T. Dauxois published in the *Bulletin de l'Union des Professeurs de Physique et de Chimie* (Nov. 2004)

The properties of several important instances of fluids found in nature—in particular their mass density ρ —depend on the altitude/depth z (oriented upwards): these fluids are said to be *stratified*. In the example of ocean water, ρ depends on depth “directly”, i.e. because it is a function of pressure \mathcal{P} which depends itself on z , but also “indirectly”, inasmuch as depth influences the salinity¹ [concentration in salt(s)], which in turn affects ρ .

The purpose of this exercise is to investigate internal waves in a stratified fluid at rest and in particular to exemplify a rather unusual dispersive behavior. Throughout, we consider a two-dimensional problem; as in the lecture, “equilibrium” quantities, related to the unperturbed fluid in absence of wave, are denoted with a subscript 0.

i. Brunt–Väisälä frequency

If a fluid particle is displaced vertically from its equilibrium position z_0 by an amount δz quickly enough, it will evolve adiabatically and without adjusting its salinity, so that when it is at $z_0 + \delta z$, its mass density ρ' differs from the equilibrium mass density $\rho_0(z_0 + \delta z)$ at that depth.

a) Considering the forces acting on the displaced fluid particle, show that Newton’s second law gives

$$\rho' \frac{d^2 \delta z}{dt^2} = -g[\rho' - \rho_0(z_0 + \delta z)], \quad (1)$$

with g the acceleration due to gravity.

The *Boussinesq approximation*, which will also be used in question **ii.**, consists in approximating the mass density in the inertial term [left hand side of Eq. (1)] by the equilibrium value $\rho_0(z_0)$, while still keeping the “exact” value (here ρ') in the force term.

b) For the right-hand side of Eq. (1), one introduces a “potential density”² $\bar{\rho}$ —which equals the mass density ρ under the same conditions of temperature and salinity, yet at a fixed reference pressure—such that the difference in the term between square brackets can be recast as $-(d\bar{\rho}/dz)\delta z$.

Under which condition on the derivative $d\bar{\rho}/dz$ is the equilibrium of the stratified fluid stable? In that case, what is the motion of the fluid particle? You may find it interesting to introduce the *Brunt–Väisälä “frequency”* defined by the relation $\omega_{B-V}^2 \equiv -(g/\rho_0) d\bar{\rho}/dz$.

ii. Propagation of internal waves

Starting from a state of (stratified) rest, we consider small perturbations $\delta\rho(t, x, z)$, $\delta\mathcal{P}(t, x, z)$, $\delta\vec{v}(t, x, z)$, assuming that the resulting flow is incompressible. We shall assume that the conservation of the potential density along streamlines reads

$$\frac{\partial \delta\rho}{\partial t} + \delta v_z \frac{d\bar{\rho}}{dz} = 0. \quad (2)$$

¹Bonus question to the former “Non-equilibrium physics” students: how? (neglecting the temperature variations)

²For more information, see https://en.wikipedia.org/wiki/Potential_density or http://oceanworld.tamu.edu/resources/ocng_textbook/chapter06/chapter06_05.htm.

- a) Write down the (kinematic) incompressibility condition, which will be hereafter referred to as Eq. (3).
 b) Show that the Euler equation in the Boussinesq approximation introduced above gives you to leading order the usual fundamental equation of hydrostatics, while the subleading order yields

$$\rho_0 \frac{\partial \delta v_x}{\partial t} = -\frac{\partial \delta \mathcal{P}}{\partial x} \quad , \quad \rho_0 \frac{\partial \delta v_z}{\partial t} = -\frac{\partial \delta \mathcal{P}}{\partial z} - \delta \rho g. \quad (4,5)$$

Together with Eqs. (2) and (3), you now have four equations for the four unknown fields $\delta \rho$, $\delta \mathcal{P}$, δv_x , and δv_z .

- c) Neglecting the spatial variations of ρ_0 —which seems at first paradoxical, but amounts to assuming that the typical length scale of variations $|g/\omega_{B-V}^2|$ is much larger than the vertical wavelength of waves—show that equations (2)–(5) lead to the differential equation

$$\frac{\partial^4 \delta v_z}{\partial t^2 \partial x^2} + \frac{\partial^4 \delta v_z}{\partial t^2 \partial z^2} = -\omega_{B-V}^2 \frac{\partial^2 \delta v_z}{\partial x^2}. \quad (6)$$

[Hint: consider (2) and (5) on the one hand, (3) and (4) on the other hand]

Convince yourself that δv_x , $\delta \mathcal{P}$ and $\delta \rho$ obey similar equations.

- d) Show that the harmonic ansatz $\delta v_z = \widetilde{\delta v}_z e^{-i(\omega t - \vec{k} \cdot \vec{r})}$ leads to the dispersion relation

$$\omega = \pm \omega_{B-V} \sin \beta_{\vec{k}}, \quad (7)$$

where $\beta_{\vec{k}}$ is the angle between the wave vector \vec{k} and the z -direction. What do you find surprising here? Discuss the physics for various values of ω (consider 4 cases!).

- e) Compute the phase velocity $\vec{c}_\varphi(\vec{k})$ (remember that it is directed along \vec{k}) and the group velocity $\vec{c}_g(\vec{k}) \equiv d\omega/d\vec{k}$ following from the dispersion relation (7), and compare them with each other. Your result begs for comments!

24. Energy-momentum tensor

Let \mathcal{R} denote a fixed reference frame. Show with the help of a Lorentz transformation that the Minkowski coordinates of the (local) energy-momentum tensor of a perfect fluid whose local rest frame moves with velocity \vec{v} with respect to \mathcal{R} , are given to order $\mathcal{O}(|\vec{v}|/c)$ by

$$T^{\mu\nu} = \begin{pmatrix} \epsilon & (\epsilon + \mathcal{P}) \frac{v^1}{c} & (\epsilon + \mathcal{P}) \frac{v^2}{c} & (\epsilon + \mathcal{P}) \frac{v^3}{c} \\ (\epsilon + \mathcal{P}) \frac{v^1}{c} & \mathcal{P} & 0 & 0 \\ (\epsilon + \mathcal{P}) \frac{v^2}{c} & 0 & \mathcal{P} & 0 \\ (\epsilon + \mathcal{P}) \frac{v^3}{c} & 0 & 0 & \mathcal{P} \end{pmatrix}.$$

Check the compatibility of this result with the general formula for $T^{\mu\nu}$ given in the lecture.

25. Isentropic flow

Let s resp. n denote the entropy density resp. particle number density, and u^μ the flow velocity. The entropy per particle is defined as $s/n = S/N$. Show that in an isentropic flow [$d_\mu(su^\mu) = 0$] the entropy per particle is conserved, i.e. $d(s/n)/dt = 0$.

Hint: If the covariant derivatives d_μ upset you, choose Minkowski coordinates, in which $d_\mu = \partial_\mu$.