

Fundamental equations of nonrelativistic fluid dynamics

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- Reynolds' transport theorem

- Equations of motion of a perfect fluid

May 11 & 18

- Equations of motion of a Newtonian fluid

June 8 & 15

Throughout these slides, Cartesian coordinates ($\{x_j\}$, $\{v_j\}$, ...) are used:

$$x_j = x^j, v_j = v^j \dots \text{ for } j = 1, 2, 3$$

Reynolds' transport theorem

A simple issue:

🌐 In the Eulerian description of a fluid motion, the matter contained inside a geometrical volume ("control volume") is constantly changing

👉 open system

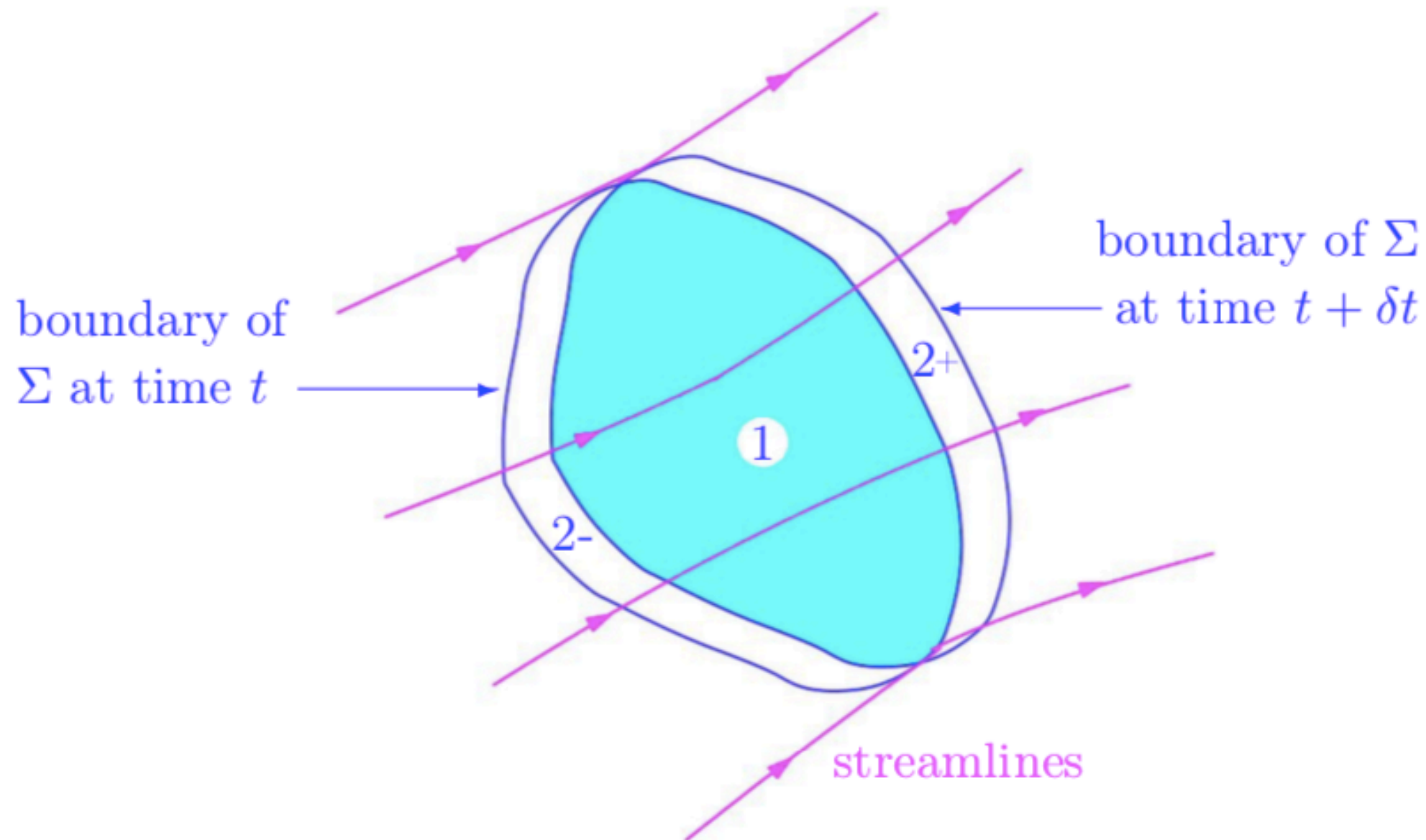
🌐 The classical equations describing the motion of mechanical systems are usually formulated for **closed systems** that are followed in their evolution

👉 Newton's 2nd law, conservation of energy & of particle number...

🌐 How do we derive equations implementing the key physics principles in terms of local "Eulerian" fields?

Reynolds' transport theorem

Idea: follow the system that occupies a control volume \mathcal{V} at time t in its motion between t and $t + \delta t$:



Reynolds' transport theorem

Idea: follow the system that occupies a control volume \mathcal{V} at time t in its motion between t and $t + \delta t$.

👉 rate of change of an extensive (physical!) quantity \mathcal{G} :

$$\frac{D\mathcal{G}(t)}{Dt} = \int_{\mathcal{V}} \frac{\partial}{\partial t} [g_m(t, \vec{r}) \rho(t, \vec{r})] d^3\vec{r} + \oint_{\partial\mathcal{V}} [g_m(t, \vec{r}) \rho(t, \vec{r}) \vec{v}(t, \vec{r})] \cdot d^2\vec{S}$$

where $g_m(t, \vec{r})$ is the local amount of \mathcal{G} per unit mass.

- 1st term: **local** rate of change (due to non-stationarity) inside \mathcal{V} ;
- 2nd term: flux of \mathcal{G} through the surface $\partial\mathcal{V}$: **convective** term.

Conservation of mass

● Global formulation:
$$\frac{DM(t)}{Dt} = 0$$

● Local formulation:

Reynolds' transport theorem gives: (specific density of mass = 1!)

$$\frac{DM(t)}{Dt} = \int_{\mathcal{V}} \frac{\partial \rho(t, \vec{r})}{\partial t} d^3 \vec{r} + \oint_{\partial \mathcal{V}} [\rho(t, \vec{r}) \vec{v}(t, \vec{r})] \cdot d^2 \vec{\mathcal{S}}$$

Gauss theorem \rightarrow

$$= \int_{\mathcal{V}} \left[\frac{\partial \rho(t, \vec{r})}{\partial t} + \vec{\nabla} \cdot [\rho(t, \vec{r}) \vec{v}(t, \vec{r})] \right] d^3 \vec{r}$$

Vanishes for every (control) volume \mathcal{V} \Rightarrow "continuity equation"

$$\frac{\partial \rho(t, \vec{r})}{\partial t} + \vec{\nabla} \cdot [\rho(t, \vec{r}) \vec{v}(t, \vec{r})] = 0$$

Newton's second law (momentum law)

● Global formulation:
$$\frac{D\vec{P}(t)}{Dt} = \vec{F}(t)$$

● Local formulation:

Reynolds' transport theorem: specific density of \vec{P} : $\vec{v}(\vec{r})$

$$\frac{D\vec{P}(t)}{Dt} = \frac{d}{dt} \left[\int_{\mathcal{V}} \vec{v}(t, \vec{r}) \rho(t, \vec{r}) d^3\vec{r} \right] + \oint_{\partial\mathcal{V}} \vec{v}(t, \vec{r}) \rho(t, \vec{r}) \vec{v}(t, \vec{r}) \cdot d^2\vec{S}$$

$$= \int_{\mathcal{V}} \rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] d^3\vec{r}$$

Using Gauss' theorem
& the continuity equation

$$\frac{D\vec{v}(t, \vec{r})}{Dt}$$

Newton's second law (momentum law)

● Global formulation:
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● Local formulation:

Reynolds' transport theorem: specific density of \vec{P} : $\vec{v}(\vec{r})$

$$\frac{D\vec{P}(t)}{Dt} = \int_{\mathcal{V}} \rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] d^3 \vec{r}$$

Forces:
$$\vec{F}(t) = \int_{\mathcal{V}} \underbrace{\vec{f}_V(t, \vec{r})}_{\text{Volume forces: external (gravity...)}} d^3 \vec{r} + \oint_{\partial \mathcal{V}} \underbrace{\vec{T}_s(t, \vec{r})}_{\text{Surface forces: inner to the fluid}} d^2 \mathcal{S}$$

Volume forces:
external (gravity...)

Surface forces:
inner to the fluid
👉 model of the fluid

A fluid model: the perfect fluid

A **perfect fluid** is a fluid in which there are no shear stresses nor heat conduction.

Does this exist? We'll discuss that again when studying non-perfect fluids, to see which properties (= "transport coefficients") have to be set to zero...

A fluid model: the perfect fluid

A perfect fluid is a fluid in which there are no shear stresses nor heat conduction.

No shear stress:

$$\vec{T}_s(t, \vec{r}) = -\underbrace{\mathcal{P}(t, \vec{r})}_{\text{pressure}} \vec{e}_n(\vec{r})$$

leading to

$$\oint_{\partial\mathcal{V}} \vec{T}_s(t, \vec{r}) d^2\mathcal{S} = -\oint_{\partial\mathcal{V}} \mathcal{P}(t, \vec{r}) \vec{e}_n(\vec{r}) d^2\mathcal{S} \stackrel{\text{divergence theorem}}{=} -\int_{\mathcal{V}} \vec{\nabla} \mathcal{P}(t, \vec{r}) d^3\vec{r}$$

Momentum law for a perfect fluid

● Global formulation:
$$\frac{D\vec{P}(t)}{Dt} = \vec{F}(t)$$

● Local formulation:

Reynolds' transport theorem: specific density of \vec{P} : $\vec{v}(\vec{r})$

$$\frac{D\vec{P}(t)}{Dt} = \int_{\mathcal{V}} \rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] d^3\vec{r}$$

Forces:

$$\vec{F}(t) = \int_{\mathcal{V}} \vec{f}_V(t, \vec{r}) d^3\vec{r} - \int_{\mathcal{V}} \vec{\nabla} \mathcal{P}(t, \vec{r}) d^3\vec{r}$$

$$\rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] = -\vec{\nabla} \mathcal{P}(t, \vec{r}) + \vec{f}_V(t, \vec{r})$$

Euler equation

Euler equation

$$\rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] = -\vec{\nabla} \mathcal{P}(t, \vec{r}) + \vec{f}_V(t, \vec{r})$$

Partial differential equation ("PDE"):

need to specify boundary conditions

- Impermeability at a wall / an obstacle: no normal component for the relative velocity

but no condition on the tangential component

- At infinity: "user-specified" (for instance: uniform)

Euler equation

$$\rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] = -\vec{\nabla} \mathcal{P}(t, \vec{r}) + \vec{f}_V(t, \vec{r})$$

Alternative formulations:

• with the density of volume forces per unit mass $\vec{a}_V \equiv \vec{f}_V / \rho$

• with the material derivative:

$$\frac{D\vec{v}(t, \vec{r})}{Dt} = -\frac{\vec{\nabla} \mathcal{P}(t, \vec{r})}{\rho(t, \vec{r})} + \vec{a}_V(t, \vec{r})$$

simple interpretation

• with the vorticity:

$$\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + \vec{\nabla} \left[\frac{\vec{v}(t, \vec{r})^2}{2} \right] - \vec{v}(t, \vec{r}) \times \vec{\omega}(t, \vec{r}) = -\frac{1}{\rho(t, \vec{r})} \vec{\nabla} \mathcal{P}(t, \vec{r}) + \vec{a}_V(t, \vec{r})$$

will be exploited in upcoming lectures

Euler equation

$$\rho(t, \vec{r}) \left[\frac{\partial \vec{v}(t, \vec{r})}{\partial t} + [\vec{v}(t, \vec{r}) \cdot \vec{\nabla}] \vec{v}(t, \vec{r}) \right] = -\vec{\nabla} \mathcal{P}(t, \vec{r}) + \vec{f}_V(t, \vec{r})$$

Alternative formulations:

● as a balance equation:

$$\frac{\partial}{\partial t} [\rho(t, \vec{r}) v^i(t, \vec{r})] + \sum_{j=1}^3 \frac{\partial \mathbf{T}^{ij}(t, \vec{r})}{\partial x^j} = f_V^i(t, \vec{r})$$

with the momentum flux tensor (= how much p^i flows in direction j)

$$\mathbf{T}^{ij}(t, \vec{r}) \equiv \mathcal{P}(t, \vec{r}) \delta^{ij} + \rho(t, \vec{r}) v^i(t, \vec{r}) v^j(t, \vec{r})$$

and the momentum density $\rho(t, \vec{r}) v^i(t, \vec{r})$.

time derivative of density + divergence of current = source term

Energy conservation in perfect fluids

$$\frac{\partial}{\partial t} \left[\frac{1}{2} \rho(t, \vec{r}) \vec{v}(t, \vec{r})^2 + e(t, \vec{r}) + \rho(t, \vec{r}) \Phi(t, \vec{r}) \right] + \vec{\nabla} \cdot \left\{ \left[\frac{1}{2} \rho(t, \vec{r}) \vec{v}(t, \vec{r})^2 + e(t, \vec{r}) + \mathcal{P}(t, \vec{r}) + \rho(t, \vec{r}) \Phi(t, \vec{r}) \right] \vec{v}(t, \vec{r}) \right\} = 0$$

with $e(t, \vec{r})$: internal energy density

and $\Phi(t, \vec{r})$: external potential, such that $\vec{f}_V(t, \vec{r}) = -\rho(t, \vec{r}) \vec{\nabla} \Phi(t, \vec{r})$

Work of pressure forces

Entropy conservation in perfect fluids

A by now familiar form:

$$\frac{\partial s(t, \vec{r})}{\partial t} + \vec{\nabla} \cdot [s(t, \vec{r}) \vec{v}(t, \vec{r})] = 0$$

with $s(t, \vec{r})$: entropy density