

Tutorial sheet 2

Discussion topic: What is the relation between the energy-momentum tensor of a perfect relativistic fluid and its internal-energy density, pressure, and four-velocity? How is the latter defined?

Throughout this exercise sheet, we do not denote the x -dependence of the various fields (and/or their components) for the sake of brevity. In addition, a set of Minkowski components $\{x^\mu\}$ is used for x .

3. Energy-momentum tensor

Let \mathcal{R} denote a fixed reference frame. Consider a perfect fluid whose local rest frame at a point x moves with velocity \vec{v} with respect to \mathcal{R} . Show with the help of a Lorentz transformation that the Minkowski components of the energy-momentum tensor of the fluid at x are given to order $\mathcal{O}(|\vec{v}|/c)$ by

$$T^{\mu\nu} = \begin{pmatrix} \epsilon & (\epsilon + \mathcal{P})\frac{v^1}{c} & (\epsilon + \mathcal{P})\frac{v^2}{c} & (\epsilon + \mathcal{P})\frac{v^3}{c} \\ (\epsilon + \mathcal{P})\frac{v^1}{c} & \mathcal{P} & 0 & 0 \\ (\epsilon + \mathcal{P})\frac{v^2}{c} & 0 & \mathcal{P} & 0 \\ (\epsilon + \mathcal{P})\frac{v^3}{c} & 0 & 0 & \mathcal{P} \end{pmatrix}.$$

Check the compatibility of this result with the general formula for $T^{\mu\nu}$ given in the lecture.

4. Equations of motion of a perfect relativistic fluid

In this exercise, we set $c = 1$ for simplicity. Show that the energy-momentum conservation equation for a perfect fluid is equivalent to the two equations

$$u^\mu \partial_\mu \epsilon + (\epsilon + \mathcal{P}) \partial_\mu u^\mu = 0 \quad \text{and} \quad (\epsilon + \mathcal{P}) u^\mu \partial_\mu u^\nu + \nabla^\nu \mathcal{P} = 0, \quad (1)$$

where $\nabla^\mu \equiv \Delta^{\mu\nu} \partial_\nu$ (see exercise 2.).

5. Isentropic flow

Consider a relativistic fluid with a single conserved charge. Let s resp. n denote the entropy density resp. the number density of the conserved charge, and u^μ the flow velocity.

Show that in an isentropic flow [$\partial_\mu (s u^\mu) = 0$] the “entropy per particle”, defined as s/n , is conserved, i.e. $d(s/n)/dt = 0$.