

Tutorial sheet 1

Discussion topic: What are the fundamental equations of the dynamics of a relativistic fluid?

Throughout this exercise sheet, we set $c = 1$, while the metric tensor has signature $(-, +, +, +)$.

1. Quantum number conservation

Consider a 4-current with components $N^\mu(x)$ obeying the continuity equation $\partial_\mu N^\mu(x) = 0$, where the $\{x^\mu\}$ are Minkowski coordinates. Show that the quantity $\mathcal{N} = \int N^0(x) d^3\vec{r}$ is a Lorentz scalar, by convincing yourself first that \mathcal{N} can be rewritten in the form

$$\mathcal{N} = \int_{x^0=\text{const.}} N^\mu(x) d^3\sigma_\mu, \quad (1)$$

where $d^3\sigma_\mu = \frac{1}{6}\epsilon_{\mu\nu\rho\sigma} d^3\mathcal{V}^{\nu\rho\sigma}$ is a 4-vector, with $d^3\mathcal{V}^{\nu\rho\sigma}$ the antisymmetric 4-tensor defined by

$$d^3\mathcal{V}^{012} = dx^0 dx^1 dx^2, \quad d^3\mathcal{V}^{021} = -dx^0 dx^2 dx^1, \quad \text{etc.}$$

and $\epsilon_{\mu\nu\rho\sigma}$ the totally antisymmetric Levi-Civita tensor with the convention $\epsilon_{0123} = +1$, such that $d^3\mathcal{V}^{\nu\rho\sigma}$ represents the 3-dimensional hypersurface element in Minkowski space.

Hint: Take some time to discuss the orientation in space-time of $d^3\sigma_\mu$ and to relate (as was done in the lecture on April 22) the “hypersurface integral” of Eq. (1) to your knowledge on surface integrals in Euclidean space.

2. A useful projector

In the lecture, the four-velocity of a fluid u^μ is defined as a timelike four-vector such that $u_\mu u^\mu = -1$.¹ Show that the tensor with components $\Delta^{\mu\nu} \equiv g^{\mu\nu} + u^\mu u^\nu$ defines a projector on the subspace orthogonal to the 4-velocity.

Denoting by d_μ the components of the (covariant) 4-gradient, we define $\nabla^\nu \equiv \Delta^{\mu\nu} d_\mu$. Can you see the rationale behind this notation?

¹Note that this also holds for the four-velocity of a point particle!