## 8.1 Correlation feurchions

So for we counciled correlation fearthous in imaginary time  $t = -i \tau$ ,  $0 \le \tau \le \beta$  defended through the path integral

 $(*) \qquad G(x-x') = Z^{-1} \int Q \varphi e^{iS[\varphi]} \varphi(x) \varphi(x')$ 

What is the operator expression while goes tile to their path integral representation?

For Tn > TL:

 $\langle \varphi(x_1) \varphi(x_2) \rangle = 3^{-1} tr \left( e^{-\beta H} \varphi(x_1) \varphi(x_2) \right)$   $= 2^{-1} tr \left( e^{-\beta H} e^{HT_1} \varphi(x_1) e^{-H(T_1 - C_2)} \varphi(x_2) e^{-HT_2} \right)$ Schrödligerpric

Here we can apply our usual derivation of the path integral by viscoshing many field and canonical—momentum eigenstates to obtain (X).
Therefore, in the operator representation, the operator must be (imaginary—) time ordered  $G(x-x') = \langle T \varphi(x) \varphi(x') \rangle$  where

TA(+) B(+'):= O(T-T') A(+)B(+') ± O(T'-T) B(+') A(+)
with the upper (lover) figure for bosonic (ferencemic) operators

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The imaginary-time correlators have no direct physical interpretation, except when  $\tau = \tau'$ . Then they represent Spatial equal-time correlations.

Ou can define verious real-trime Correlation functions. Ne will be that they can all be obtained from imaginary-time correlators by appropriate analytic Contienation.

For any field operator A(t), B(t) (like, e.g. A(t)=\P(\xi,\xi))

one defines the Dightman functions

 $\Delta_{AB}^{2}(t) := \langle A(t)B(0) \rangle$  $\Delta_{AB}^{2}(t) := \pm \langle B(0)A(t) \rangle$  for  $\{bosonic\}$  operators

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 $\Delta_{AB}^{ret}(t) := i\Theta(t) \langle [A(t), B(0)]_{\pm} \rangle$  referred correlator  $\Delta_{AB}^{cod}(t) := -i\Theta(-t) \langle [A(t), B(0)]_{\pm} \rangle$  referred

Before we discuss relations between these functions we'll counder an application:

8.2 Partile production

typical Lituation.

(i) photon production in quark-gluon pleisure opents & gluons interact respectly. Then a photon is produced it can most likely escape from the System

(ii) early universe: matter in thurnal equilibrium some non-Standard Model particles like sterile neutrinos, gravitius, or axian are very weakly Compled. If they are produced they propagate freely To compute probabilities we need a normalized 1- passible state:

$$|\Psi\rangle = \int \frac{d^3r}{(2\pi)^3 2E_r} \Psi(\vec{r}) |\vec{r}\rangle$$

1 1 particle with momentum of

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momenteur space wave-function

〈大け〉>=2En (2n) d(アード)

normalizeition:

$$1 = \langle \Psi | \Psi \rangle = \int \frac{d^{3} r d^{3} r'}{(2\pi)^{6} 4 E_{r} E_{r'}} \Psi(\vec{r}) \Psi(\vec{r}') \langle \vec{r} | \vec{r}' \rangle$$

$$= \int \frac{d^{3} r}{(2\pi)^{3} 2 E_{r}} | \Psi(\vec{r})|^{2} = 0$$

probability to find 1 partice with momentum is in dip:

$$\frac{dP = \frac{1}{2E_r} |\Psi(\vec{r})|^2 \frac{d^3 r}{(2\pi)^3} = \frac{1}{2E_r} |\langle \vec{r} | \Psi \rangle|^2 \frac{d^3 r}{(2\pi)^3}$$

Now we are ready to deal with particle production.

Let X be some very weakly coupled field/particle. Intial condition: no X, but otherwish thermal equilibrium

Lier = X. J , J : operator containing the thermal fields X creates X- particles

cheraction picture  $(x = \bar{x})$  like photon, Majoreure newtrino)  $\chi(x) = \int \frac{d^3r}{(p\pi)^3 2E_p} \left\{ e^{-ipx} a_p | u_p + e^{ipx} a_p v \right\}$ 

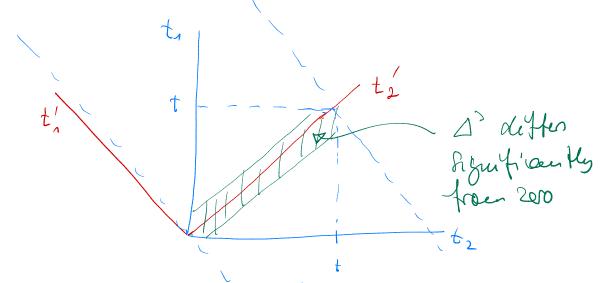
u, v: scalar, spilor of polovization vector

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time wouthhow:  $|\hat{\Psi}\rangle = H_{int} |\Psi\rangle$   $|\Psi\rangle = [n-i] \int_{0}^{t} dt' H_{int}(t') + ... ] |i|^{k} |i| |i|^{k} |i|^$ 

Sum over all f and thermally average over h'>  $\langle 1 \langle f \not h \rangle \oplus f' \rangle = \int dt \int dx_1 \int dx_2 \int dx_2 e^{i p(x_1 - x_2)} \partial v^* \langle f(x_2) f(x_3) \rangle$   $= + \Delta \int_{\mathcal{T}} f(x_1 - x_2) \quad \text{for } f' \in \text{fermionic}$  Wightman fct,  $x_1' := x_1 - x_2 \qquad \qquad t_2' := t_1 + t_2$ 

 $\left|\frac{\partial(t_1',t_2')}{\partial(t_1,t_2)}\right| = \left|\frac{1}{1} - \frac{1}{1}\right| = 2$ 



Since the planue is thermal, it "forgets" about it's pearly after a certain time. Therefore  $\Delta^2(x, -x_2)$  differ significantly from zoo only up to a certain time difference  $|t_1-t_2|$ . Thus instead of integrating over  $0 \le t_1, t_2 \le t$  we may integrate our the much larger region  $-\infty < t, < \infty$ ,  $0 \le t_2' \le 2t$  without dianging the value of the integral:  $(1 < \frac{1}{2} + 1)^2 = \pm \sqrt{\frac{1}{2}} \int dt_2' \int dt_3' \int d^3x_3' = 0 \times \Delta^2(x_3') = i + x_3$ 

$$dT = \frac{dP}{Vt} = \pm \frac{d^3p}{(2\pi)^3 2E_p} \int d^4x \quad \nabla U^* \Delta_{33}^2 + (x) e^{ipx}$$

upper to for fermions