

7.4 Non-relativistic limit

Consider system on non-relativistic electrons (mass m)
photons with energies $\ll m$

We should recover non-relativistic QM plus relativistic corrections

Let $v \ll 1$ be a typical velocity of the electrons,

idea: expand in v .

estimates: momenta $\sim mv$
kinetic energies $\sim mv^2$

So far we have used the Weyl rep. of γ -matrices

Here it is more convenient to use Dirac rep.

$$\gamma^0 = \begin{pmatrix} \mathbb{1} & 0 \\ 0 & -\mathbb{1} \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}$$

which is related to Weyl rep. by the unitary transformation

$$\gamma^M_{\text{Dirac}} = U \gamma^M_{\text{Weyl}} U^\dagger \quad \text{with} \quad U = \frac{1}{\sqrt{2}} \begin{pmatrix} \mathbb{1} & \mathbb{1} \\ -\mathbb{1} & \mathbb{1} \end{pmatrix}$$

$$\text{free Dirac field: } \psi = \sum_s \int \frac{d^3 p}{(2\pi)^3 2p^0} \left\{ e^{-ip \cdot x} u_s(\vec{p}) a_s(\vec{p}) + e^{ip \cdot x} v_s(\vec{p}) b_s^\dagger(\vec{p}) \right\}$$

$$(p - m) u = 0 \quad p = m \gamma^0 + \mathcal{O}(v)$$

$$(p - m) = m \begin{pmatrix} 0 & 0 \\ 0 & -2\mathbb{1} \end{pmatrix} \Rightarrow U = \begin{pmatrix} \xi \\ 0 \end{pmatrix} + \mathcal{O}(v)$$

lower components are small

$$p \cdot x = p^0 t - \vec{p} \cdot \vec{x}$$

$$p^0 t = \sqrt{m^2 + \vec{p}^2} t = m \left(1 + \frac{\vec{p}^2}{2m^2} + \mathcal{O}(v^4) \right) t$$

estimate $\vec{x} \sim vt \Rightarrow \vec{p} \cdot \vec{x} \sim mtv^2$

expectation: the part of ψ which is relevant here obeys

$$\psi(x) = e^{-imt} \cdot (\text{slowly varying terms})$$

slowly means $\sim mv^2$

Ansatz:

$$\psi(x) = e^{-imt} \begin{pmatrix} \chi(x) \\ \eta(x) \end{pmatrix}$$

expectation: $|\chi| \gg |\eta|$

Insert this into

$$L_e = \bar{\psi} (i\mathcal{D} - m)\psi$$

$$i\mathcal{D}_0 \psi = e^{-imt} (m + i\mathcal{D}_0) \begin{pmatrix} \chi \\ \eta \end{pmatrix}$$

$$L_e = (\chi^\dagger, \eta^\dagger) \gamma^0 (\gamma^0 m + i\gamma^0 \mathcal{D}_0 + i\vec{\gamma} \cdot \vec{\mathcal{D}} - m) \begin{pmatrix} \chi \\ \eta \end{pmatrix}$$

$$\vec{\gamma} \cdot \vec{\mathcal{D}} = \gamma^\mu \mathcal{D}_\mu, \quad \mathcal{D}_\mu = \partial_\mu + ieA^\mu$$

$$\gamma^0 \vec{\gamma} \cdot \vec{\mathcal{D}} = \gamma^0 \gamma^\mu \mathcal{D}_\mu = \begin{pmatrix} \mathbb{1} & \\ & -\mathbb{1} \end{pmatrix} \begin{pmatrix} \sigma^\mu \\ -\sigma^\mu \end{pmatrix} \mathcal{D}_\mu$$

$$= \begin{pmatrix} 0 & \sigma^\mu \\ \sigma^\mu & 0 \end{pmatrix} \mathcal{D}_\mu = \begin{pmatrix} 0 & \vec{\sigma} \cdot \vec{\mathcal{D}} \\ \vec{\sigma} \cdot \vec{\mathcal{D}} & 0 \end{pmatrix}$$

$$L_e = (x^\dagger, \eta^\dagger) \overset{Q.F.T}{\underbrace{\begin{pmatrix} iD_0 & i\vec{\sigma} \cdot \vec{D} \\ i\vec{\sigma} \cdot \vec{D} & 2m + iD_0 \end{pmatrix}}_{\begin{pmatrix} iD_0 x + i\vec{\sigma} \cdot \vec{D} \eta \\ i\vec{\sigma} \cdot \vec{D} x + (2m + iD_0) \eta \end{pmatrix}}} \begin{pmatrix} x \\ \eta \end{pmatrix}$$

$$= x^\dagger iD_0 x + x^\dagger i\vec{\sigma} \cdot \vec{D} \eta + \eta^\dagger i\vec{\sigma} \cdot \vec{D} x + \eta^\dagger (2m + iD_0) \eta$$

x "light", η "heavy"

We can integrate out η^\dagger, η in the path integral. Since the integral is Gaussian, this is equivalent to solving the eq. of motion.

$$(*) \quad i\vec{\sigma} \cdot \vec{D} x + (2m + iD_0) \eta = 0$$

(formal) solution:

$$\eta = \frac{1}{2m + iD_0} (-i) \vec{\sigma} \cdot \vec{D} x$$

assume that A_μ is sufficiently small such that $D_\mu \sim \partial_\mu$

Then $D_0 \sim m v^2$, $\vec{D} \sim m v$ and $\eta \sim v x$ as expected.

Insert (*) in L_e

$$L_e = x^\dagger iD_0 x + x^\dagger i\vec{\sigma} \cdot \vec{D} \eta \quad \Leftrightarrow$$

$$(*) \quad L_e = x^\dagger \left[iD_0 + \vec{\sigma} \cdot \vec{D} \frac{1}{2m + iD_0} \vec{\sigma} \cdot \vec{D} \right] x$$

So far: no approximations

non-relativistic limit

now we can expand in iD_0/m . At leading order

$$\frac{1}{2m + iD_0} = \frac{1}{2m}$$

$$\mathcal{L}_e = \chi^\dagger \left[iD_0 + \frac{(\vec{\sigma} \cdot \vec{D})^2}{2m} \right] \chi$$

$$\sigma^i \sigma^j = \frac{1}{2} \{ \sigma^i, \sigma^j \} + \frac{1}{2} [\sigma^i, \sigma^j] = \delta^{ij} + i \epsilon^{ijk} \sigma^k$$

$$\begin{aligned} (\vec{\sigma} \cdot \vec{D})^2 &= \sigma^i \sigma^j D_i D_j = \vec{D}^2 + i \epsilon^{ijk} \sigma^k D_i D_j \\ &= \vec{D}^2 + \frac{i}{2} \sigma^k \epsilon^{kij} [D_i, D_j] \end{aligned}$$

$$\begin{aligned} [D_i, D_j] &= (\partial_i - ie A_i)(\partial_j - ie A_j) - (i \leftrightarrow j) \\ &= -ie (\partial_i A_j - \partial_j A_i) = +ie (\partial_i A_j - \partial_j A_i) \end{aligned}$$

$$\begin{aligned} (\vec{\sigma} \cdot \vec{D})^2 &= \vec{D}^2 + i \sigma^k \epsilon^{kij} ie \partial_i A_j = \vec{D}^2 - e \vec{\sigma} \cdot \nabla \times \vec{A} \\ &= \vec{D}^2 - e \vec{\sigma} \cdot \vec{B} \end{aligned}$$

$$\mathcal{L}_e = \chi^\dagger \left[iD_0 + \frac{1}{2m} \left(\vec{D}^2 - g e \vec{B} \cdot \frac{\vec{\sigma}}{2} \right) \right] \chi$$

with

$$g = 2$$

eq. of motion

$$\left[iD_0 + \frac{1}{2m} (\vec{D}^2 - g e \vec{B} \cdot \frac{\vec{\sigma}}{2}) \right] \chi = 0$$

$$iD_0 = i\partial_t + eA_0 \quad \rightarrow$$

$$i\partial_t \chi = \left[-\frac{1}{2m} (\nabla + ie\vec{A})^2 - eA_0 + g \frac{e}{2m} \vec{B} \cdot \frac{\vec{\sigma}}{2} \right] \chi$$

Pauli equation

describes non-relativistic particle with charge (-e)

$$H = \frac{1}{2m} (\vec{p} + e\vec{A})^2 - eA_0 + g \frac{e}{2m} \vec{B} \cdot \vec{S}$$

$$\text{with } \vec{S} = \frac{\hbar}{2} \vec{\sigma}$$

energy of magnetic moment $\vec{\mu}$: $H_{int} = -\vec{\mu} \cdot \vec{B}$

for orbital motion of a particle with charge q:

$$\vec{\mu} = \frac{1}{2} q \vec{r} \times \vec{v} = +\frac{1}{2} \frac{q}{m} \vec{r} \times \vec{p} = \frac{q}{2m} \vec{L}$$

for the orbital motion one would find $g = 1$

Our result $g = 2$ for the spin is quite remarkable, and can't be understood classically!

It is not exact, there are QED corrections:

$$\frac{\hbar}{2m} \vec{\sigma} \cdot \vec{\nabla} + \frac{\hbar}{2m} \vec{\sigma} \cdot \vec{\nabla} \times \vec{A} \quad (\text{with } \vec{p} \gg m) \quad + \text{higher orders.}$$

Expanding (*) to next order gives relativistic corrections: Darwin term, spin-orbit coupling